



Nuclear Physics Underground and Beyond: LUNA Winter meeting
2026, Jan. 28 - 30, Debrecen, Hungary



Progress of Jinping Underground Nuclear Astrophysics experiment (JUNA)

Taoyu Jiao for the JUNA collaboration

HUN-REN Konkoly Observatory

Thanks NSFC, Yalong power, THU, CAS and CNNC



Nuclear Astrophysics in China



JUNA experiment

Direct in Gamow window
(underground)

RIBLL and LEAF experiment

Direct in higher energy

CIAE experiment

In-direct measurements

BRIF, CSR experiment

Nuclear decay

CSR experiment

Nuclear mass

theoretical study

Nuclear astrophysics and sensitivity study

theoretical study

Shell model and mean field calculation

theoretical study

Shell model and mean field calculation

theoretical study

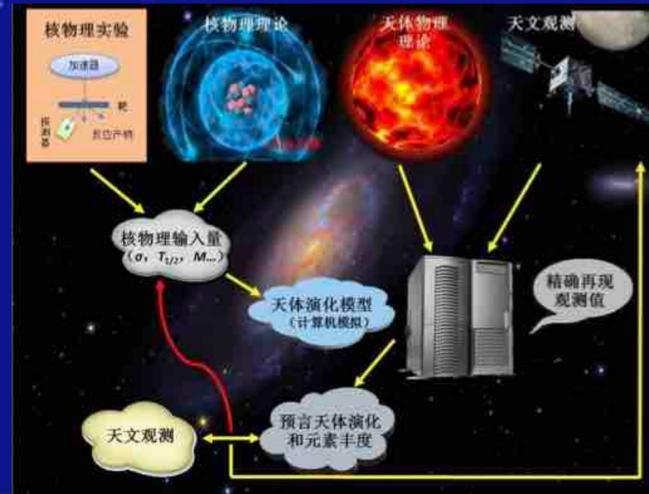
Nuclear astrophysics network calculation

Reaction rate database

Nuclear input database

Mass data compilation in China

Mass and decay rate database

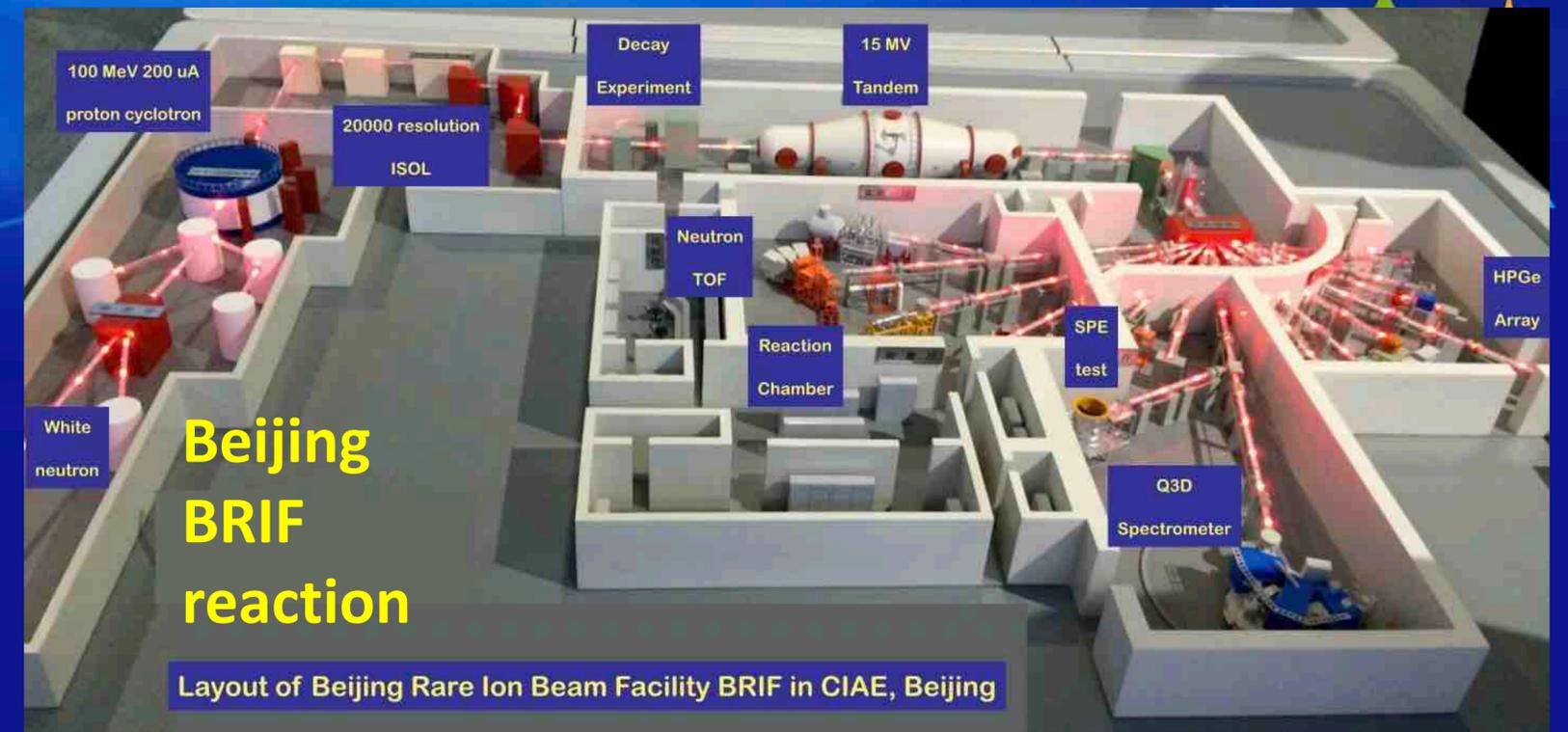


Major facilities in China

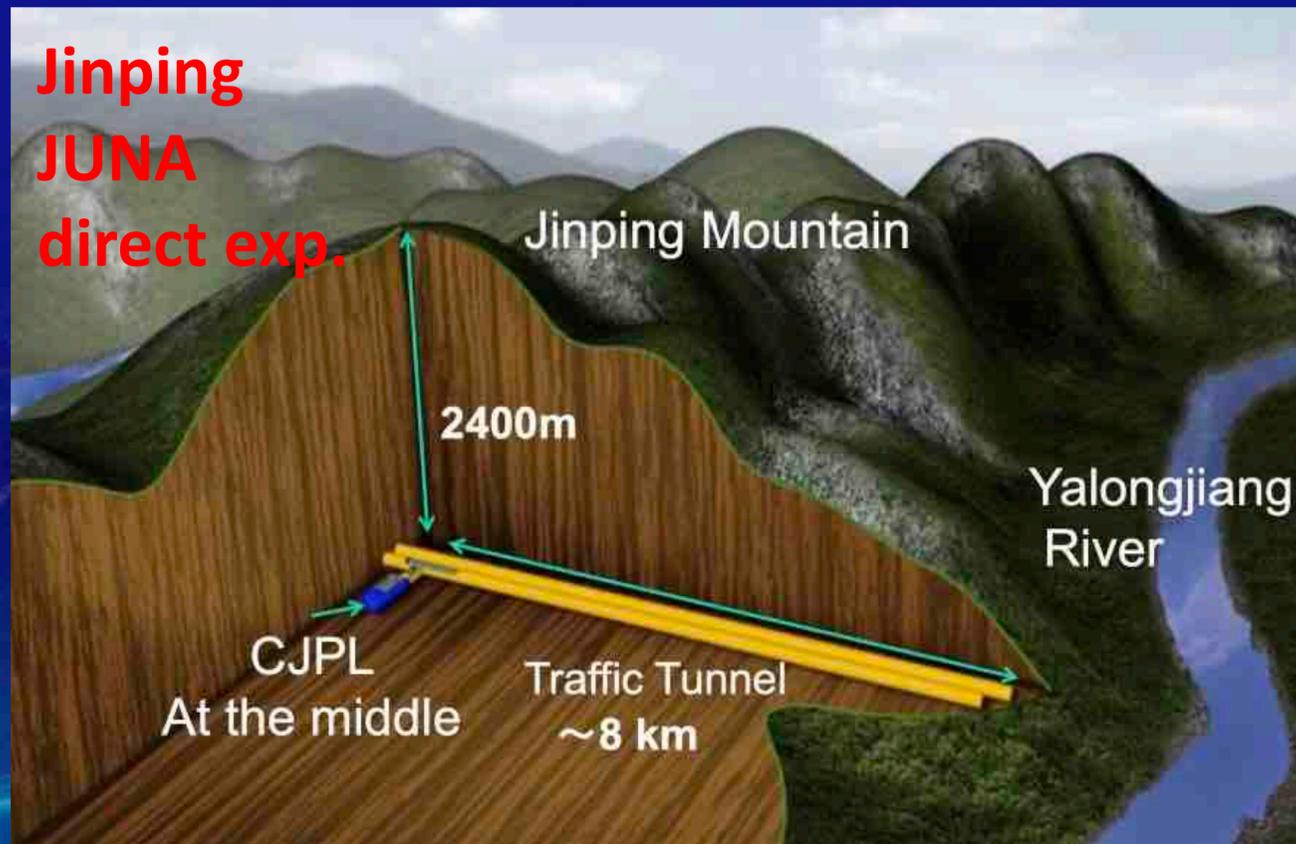
LAMOST
observation



FAST
observation



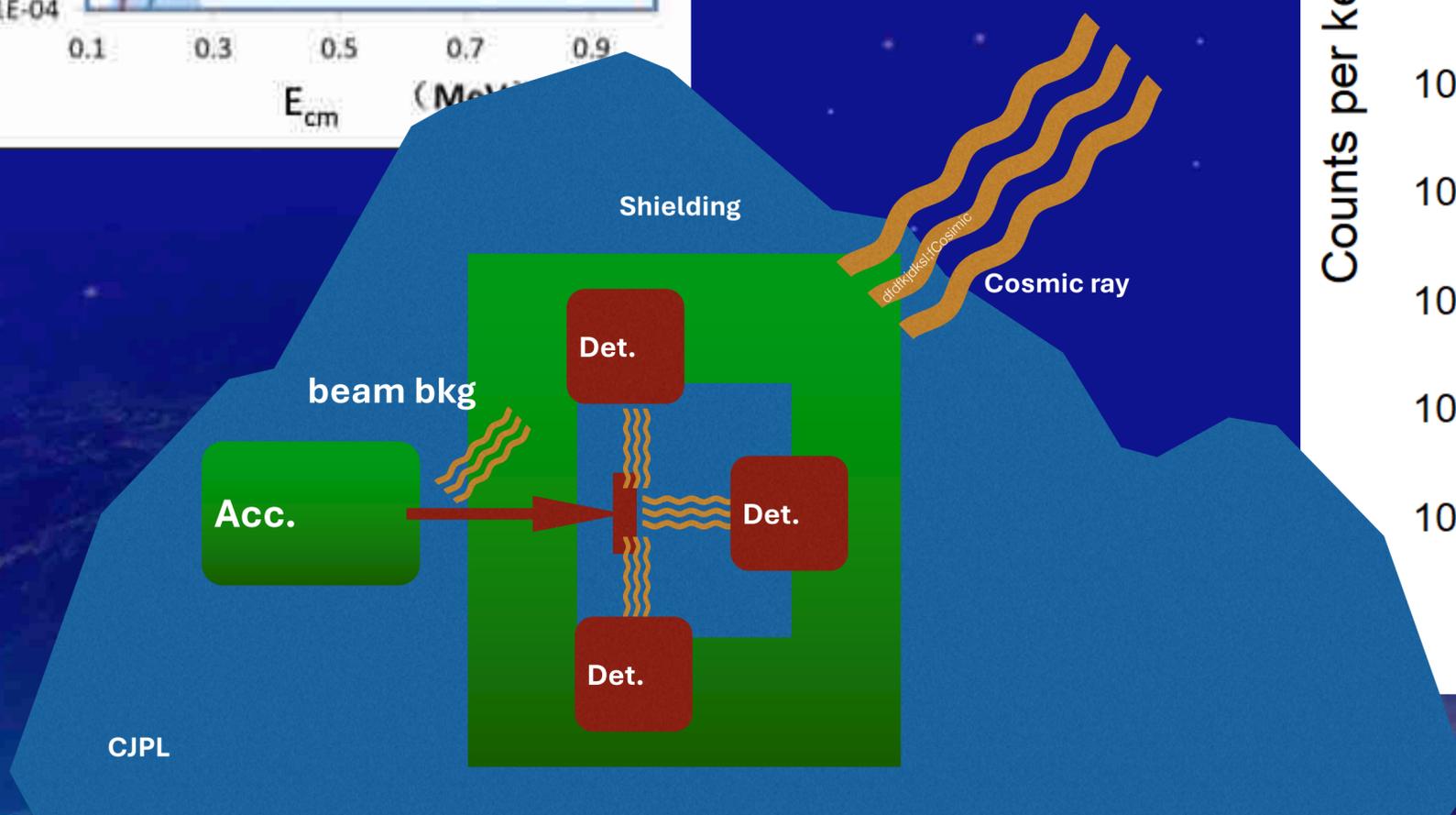
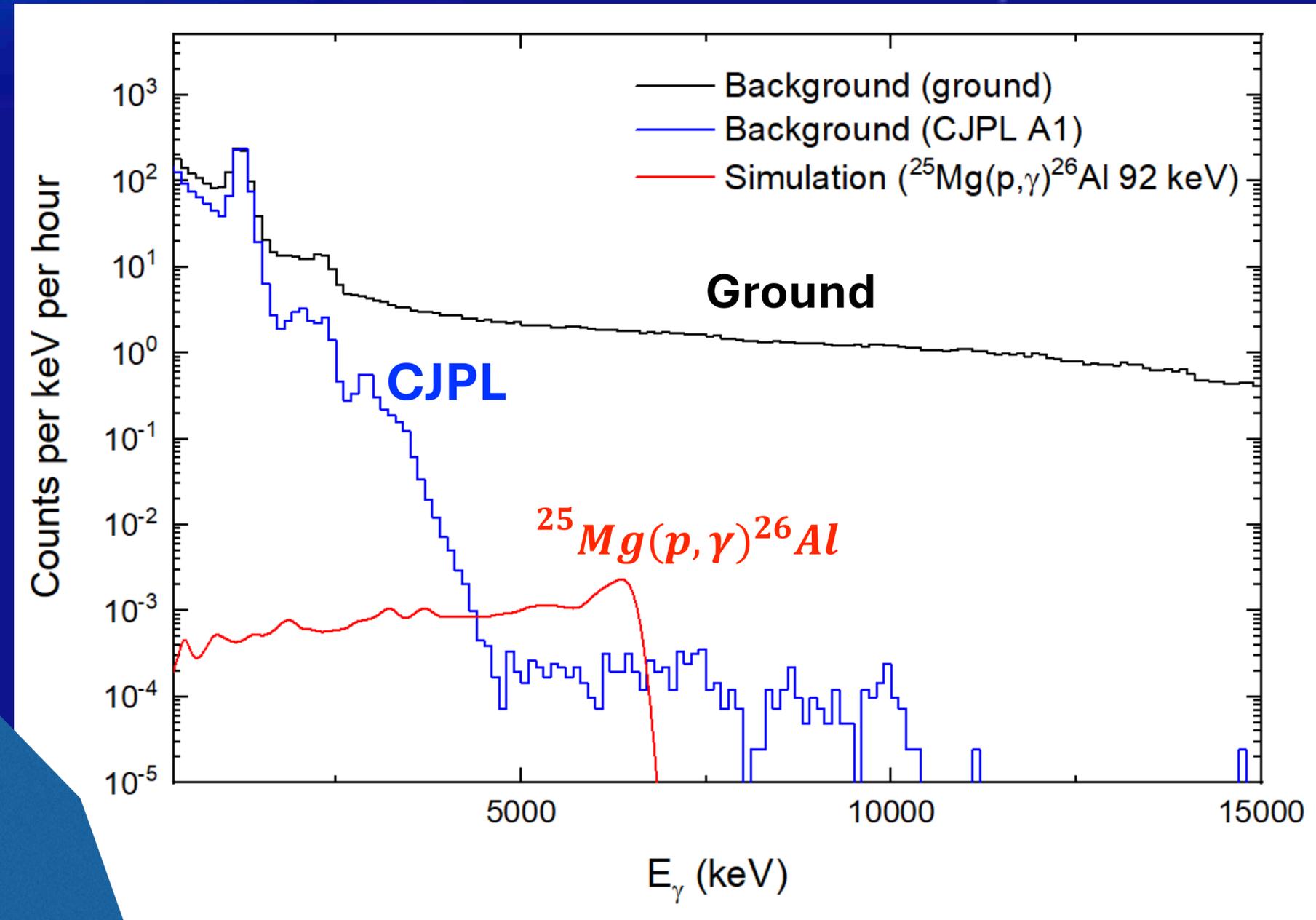
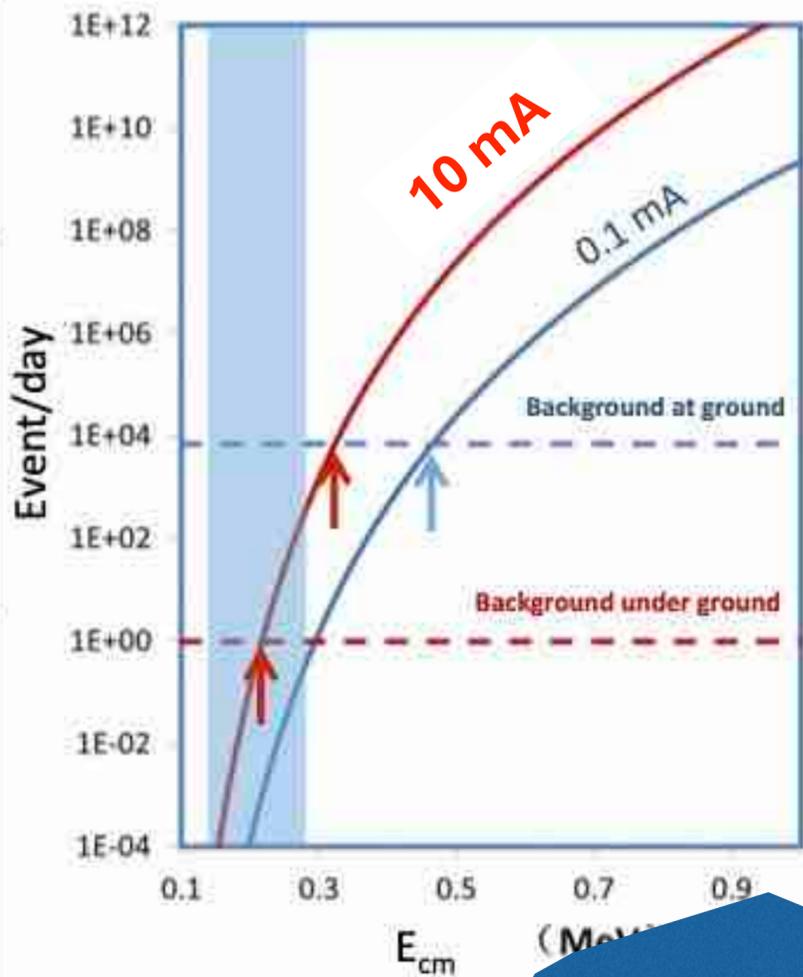
Jinping
JUNA
direct exp.



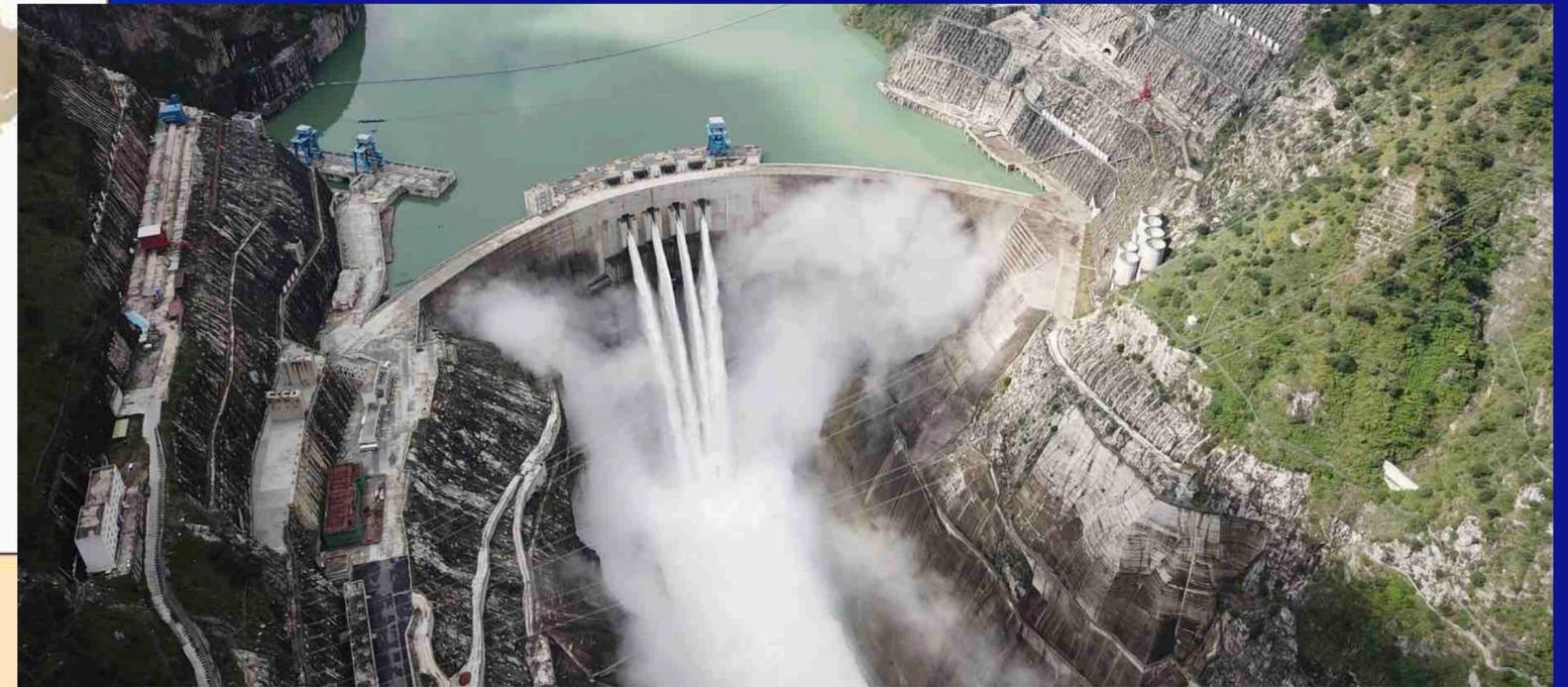
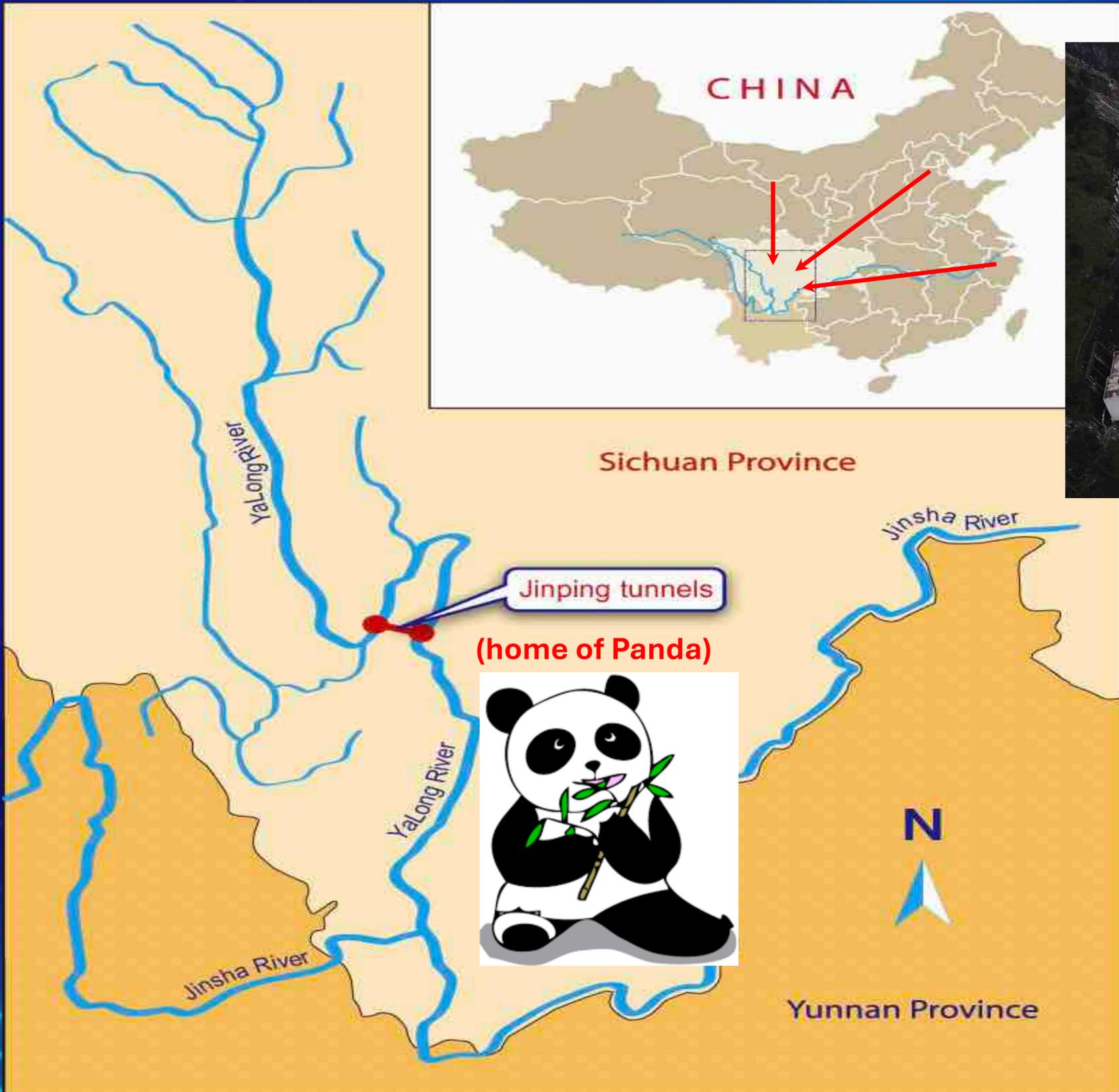
Lanzhou
CSR
mass, decay



Low bkg + high intensity



China Jinping Underground Laboratory(CJPL)

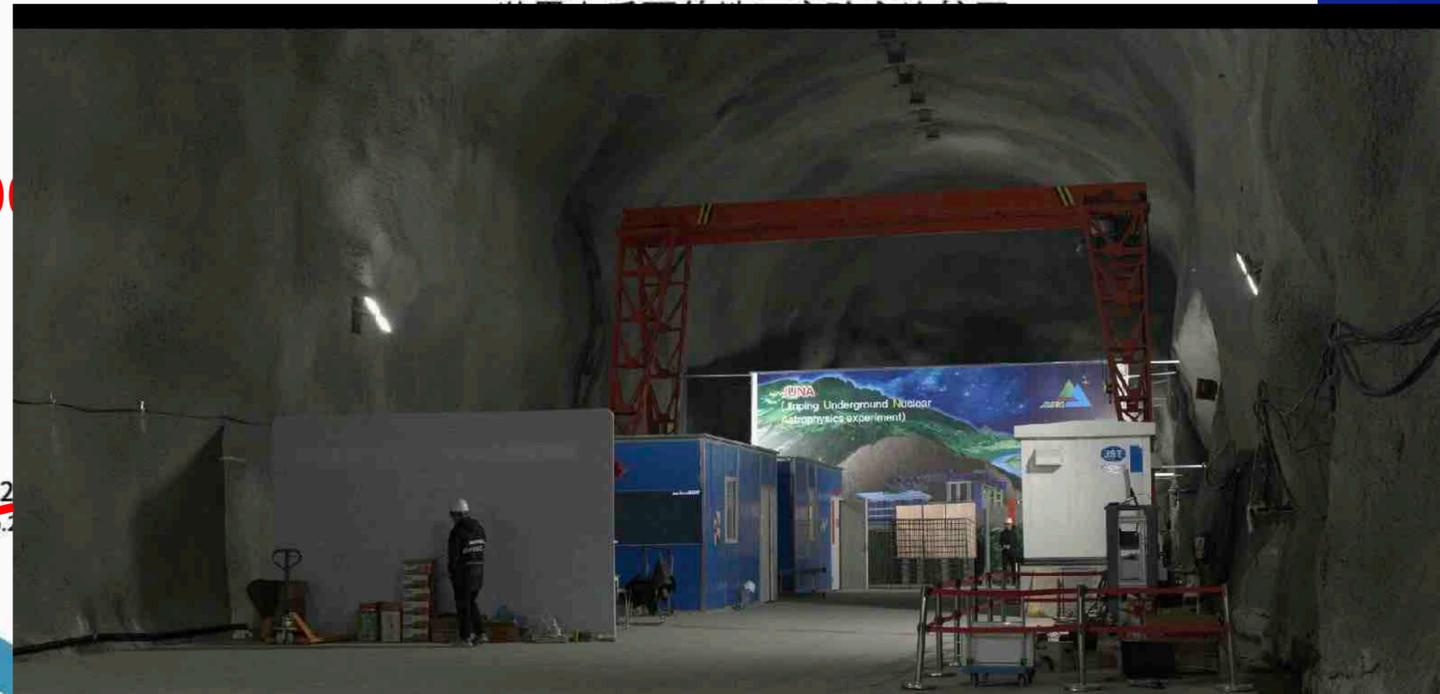






CJPL-II

Space: 300,000



JUNA dream team

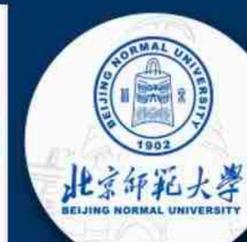
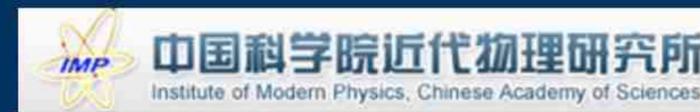


Group leader



Weiping Liu
 $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$

Yangping Shen, CIAE
Jun Su, BNU
PI



Acc. installation
Arjun Li

A1
construction
Hongwei
Yang



Site support
Xiaopan Cheng



Acc. operation
Long Zhang



Bing Guo
 $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$
CIAE



Xiaodong Tang
 $^{13}\text{C}(\alpha, n)^{16}\text{O}$
Ion source IMP



Zhihong Li
 $^{25}\text{Mg}(p, \gamma)^{26}\text{Al}$
CIAE
Jun Su, BNU



Jianjun He
 $^{19}\text{F}(p, \alpha)^{16}\text{O}$
BNU



Gang Lian
Lab. exp. sup.
CIAE



Shuo Wang
 $^{14}\text{N}(p, \gamma)^{15}\text{O}$
SDU

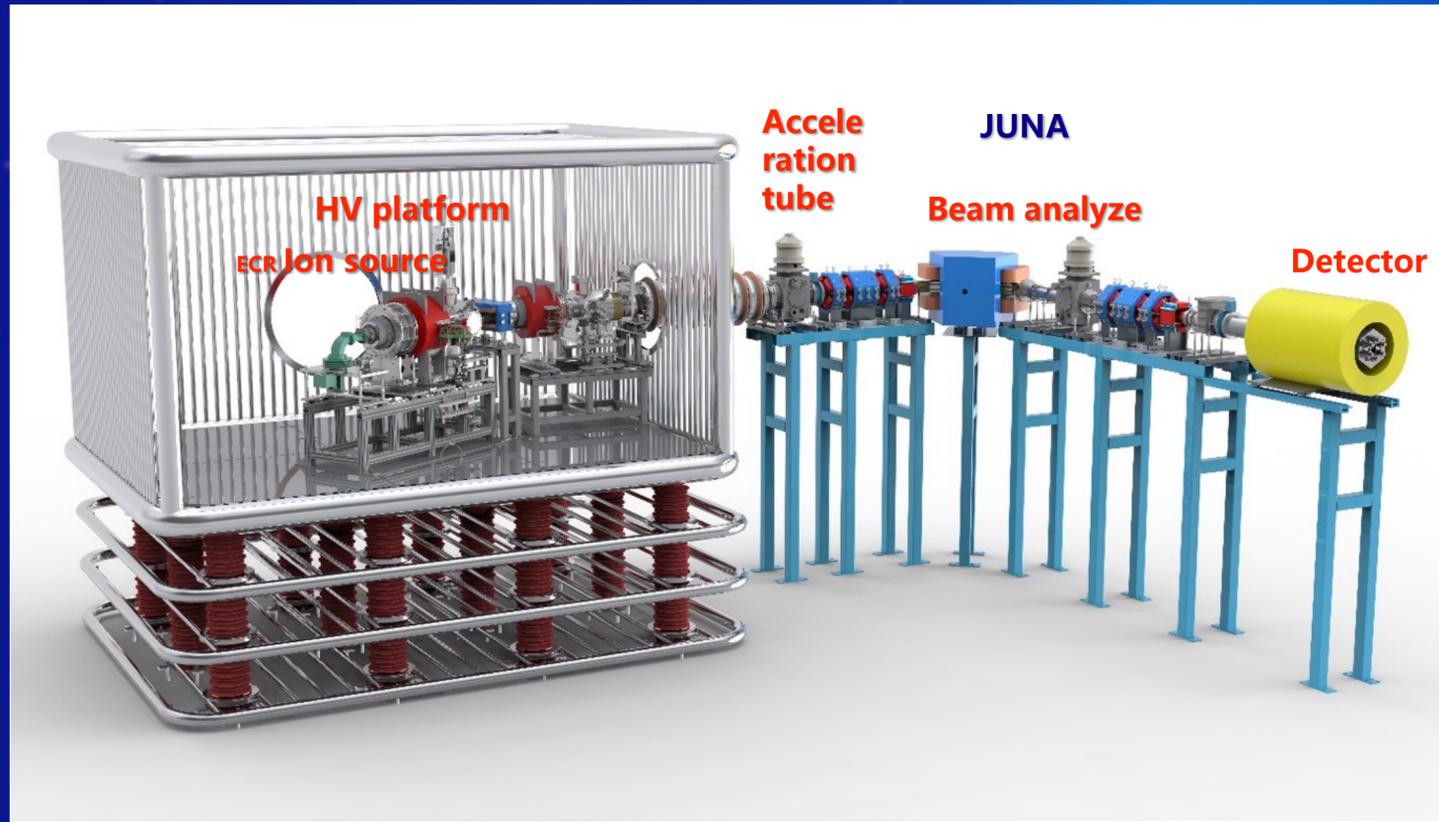


Bao
Quncui, CIAE
Liangting
Sun, IMP
Ion source
and acc.

Supported by the National Natural Science Foundation of China, Grant No. 11490560, 2015

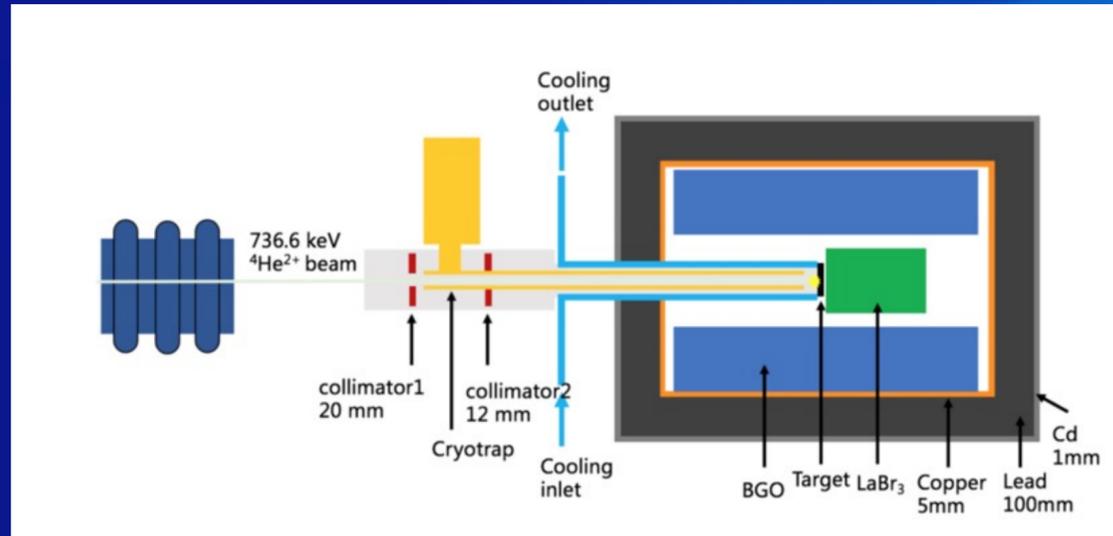
WPL et al., Sci. China 59(2016)2

Capabilities

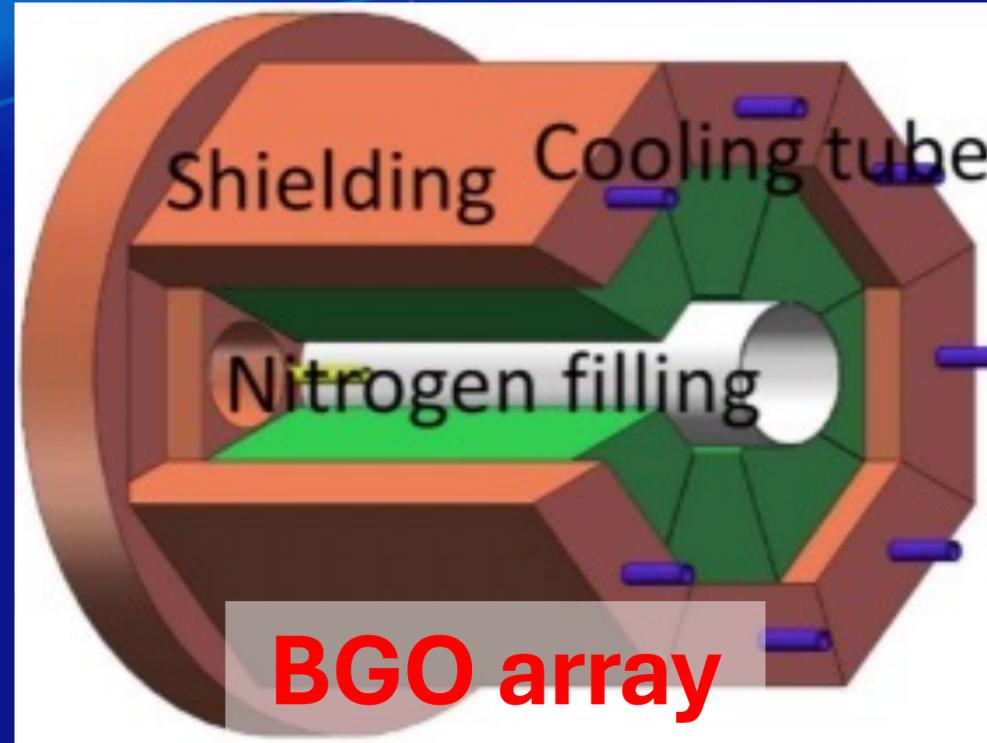


	cosmic muon bkg ($\text{cm}^{-2} \text{s}^{-1}$)	beam energy (keV)			beam intensity (emA)			energy stability
		H^+	He^+	He^{2+}	H^+	He^+	He^{2+}	
JUNA	2×10^{-10}	50-400	50-400	100-800	10	10	2	0.04%

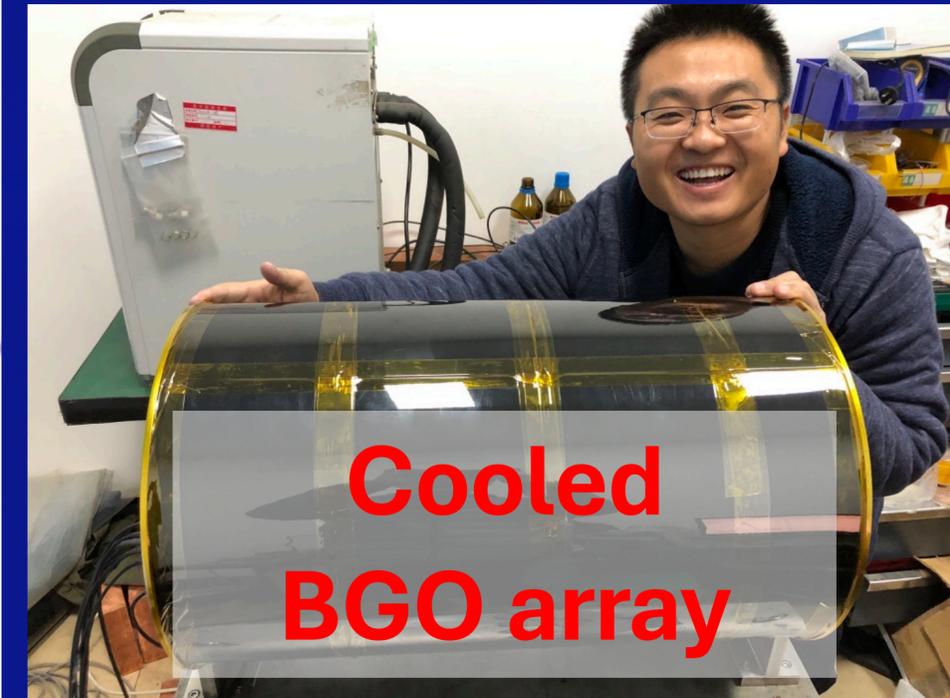
Detection systems



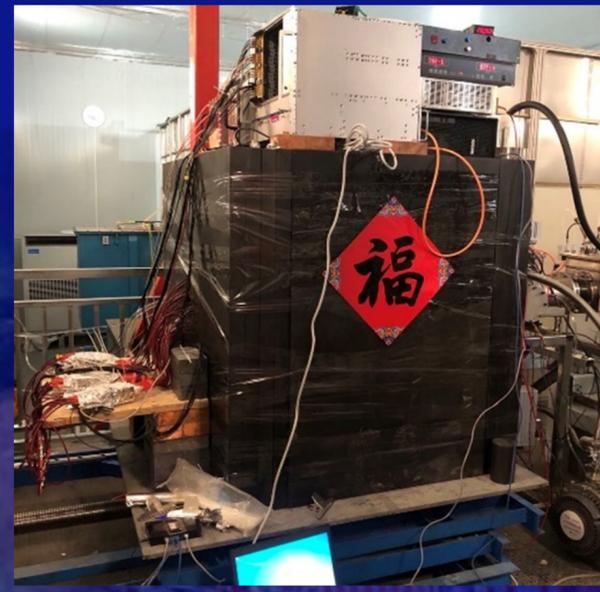
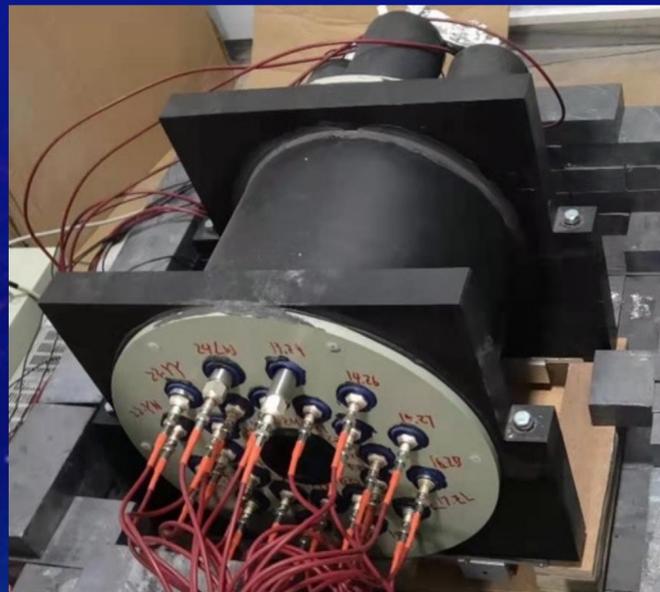
LaBr₃+BGO array



BGO array



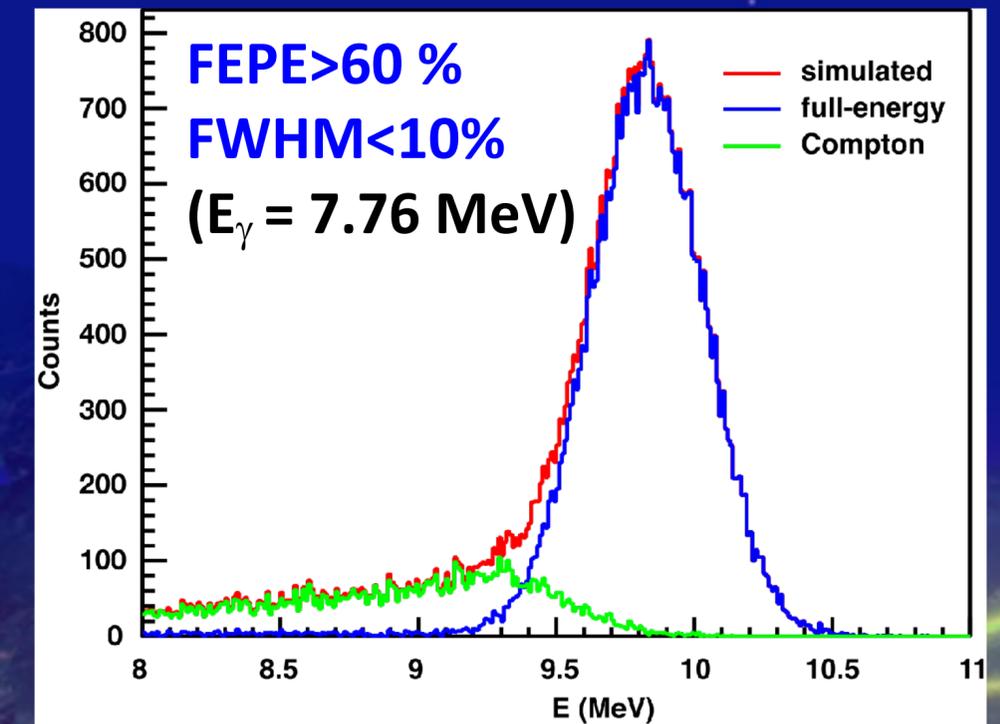
Cooled BGO array



³He detector array
26%@2.5MeV neutron

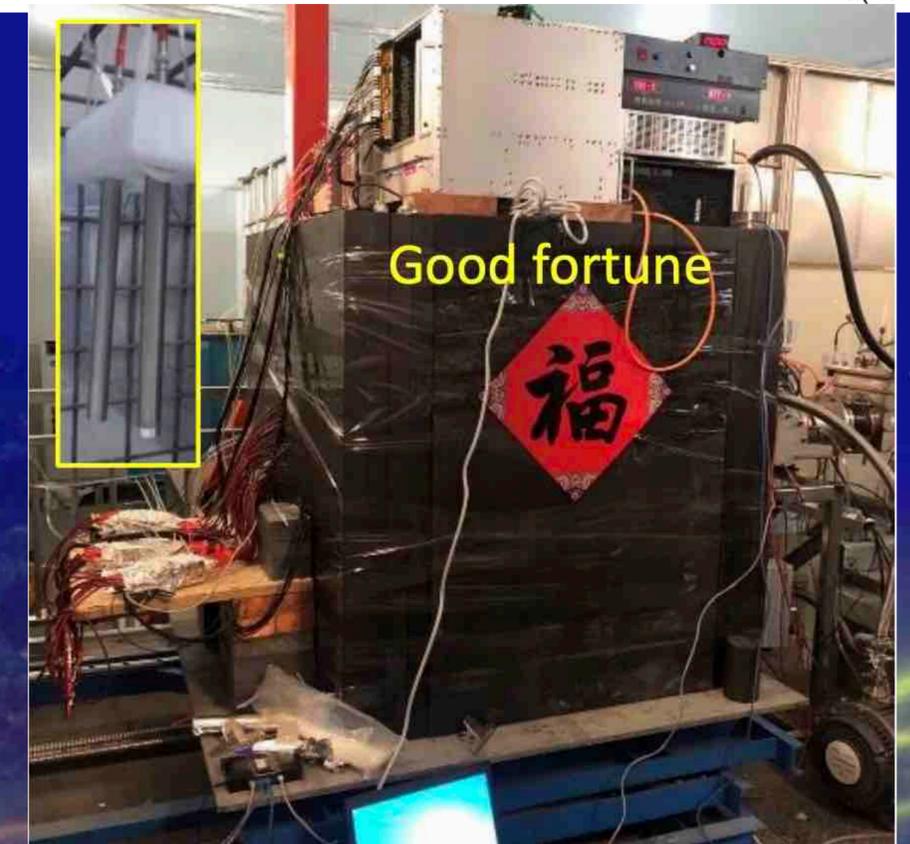
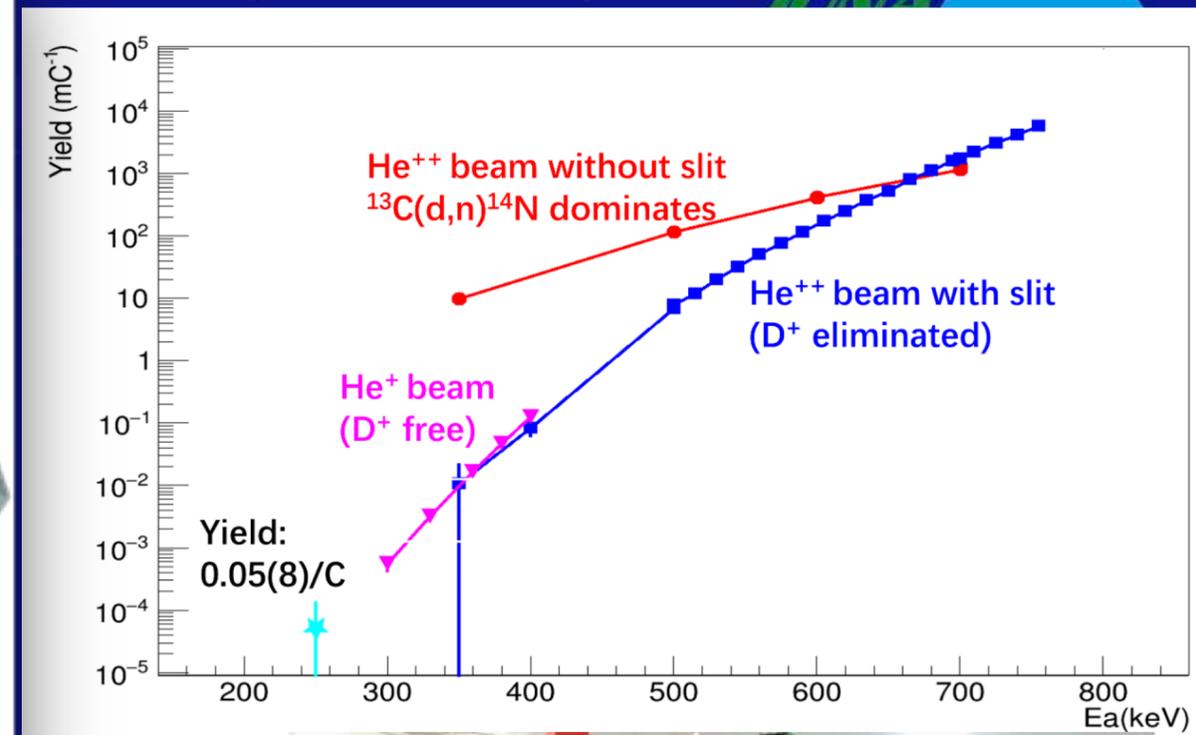
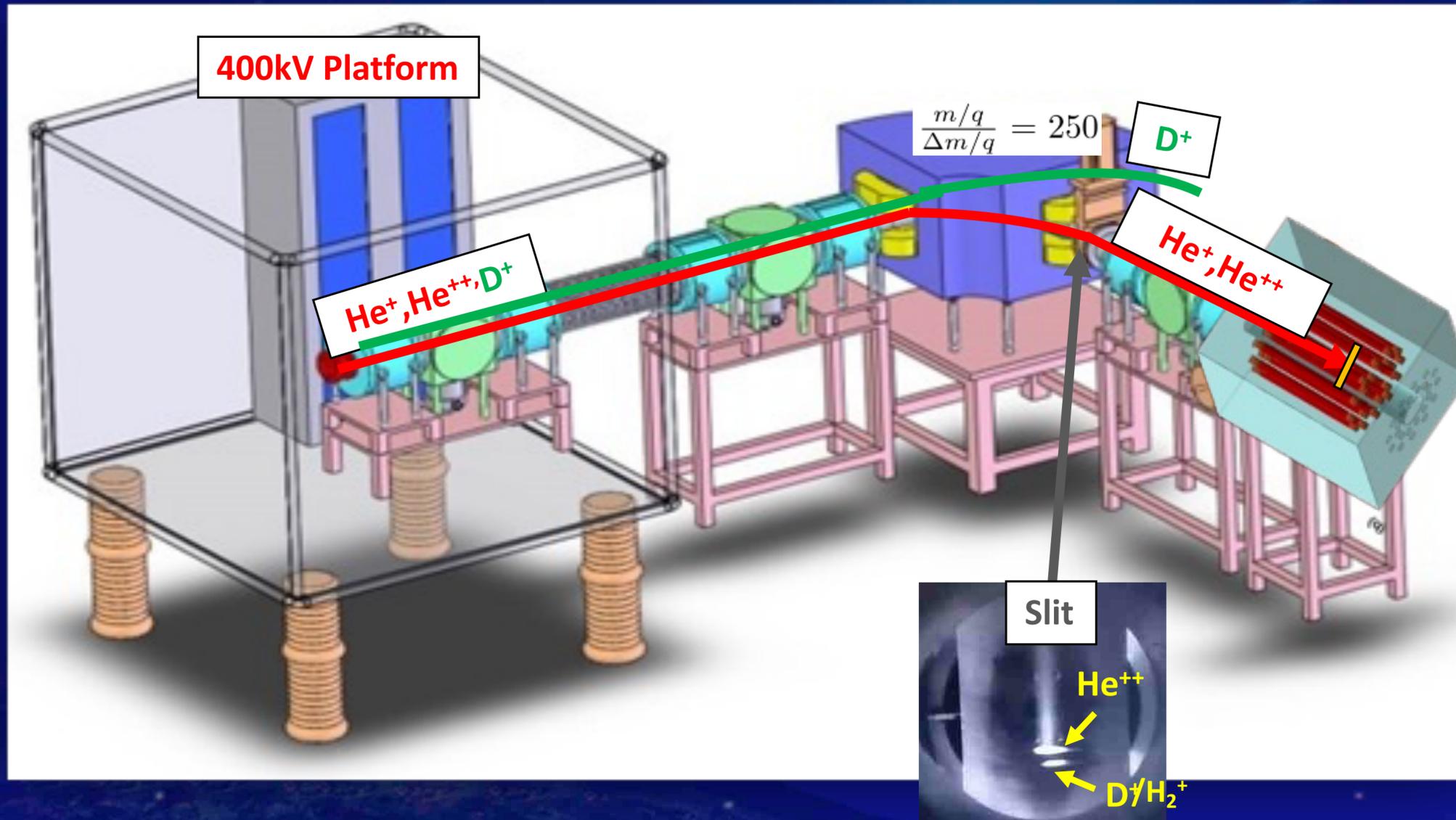


Silicon array



WPL et al., Sci. China 59(2016)2

$^{13}\text{C}(\alpha,n)^{16}\text{O}$ Measurement



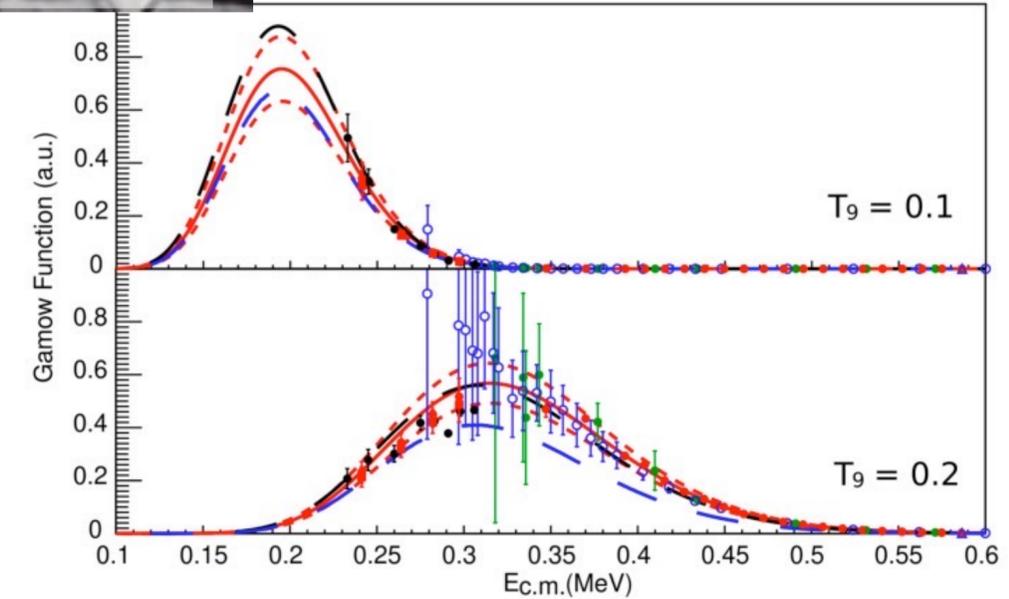
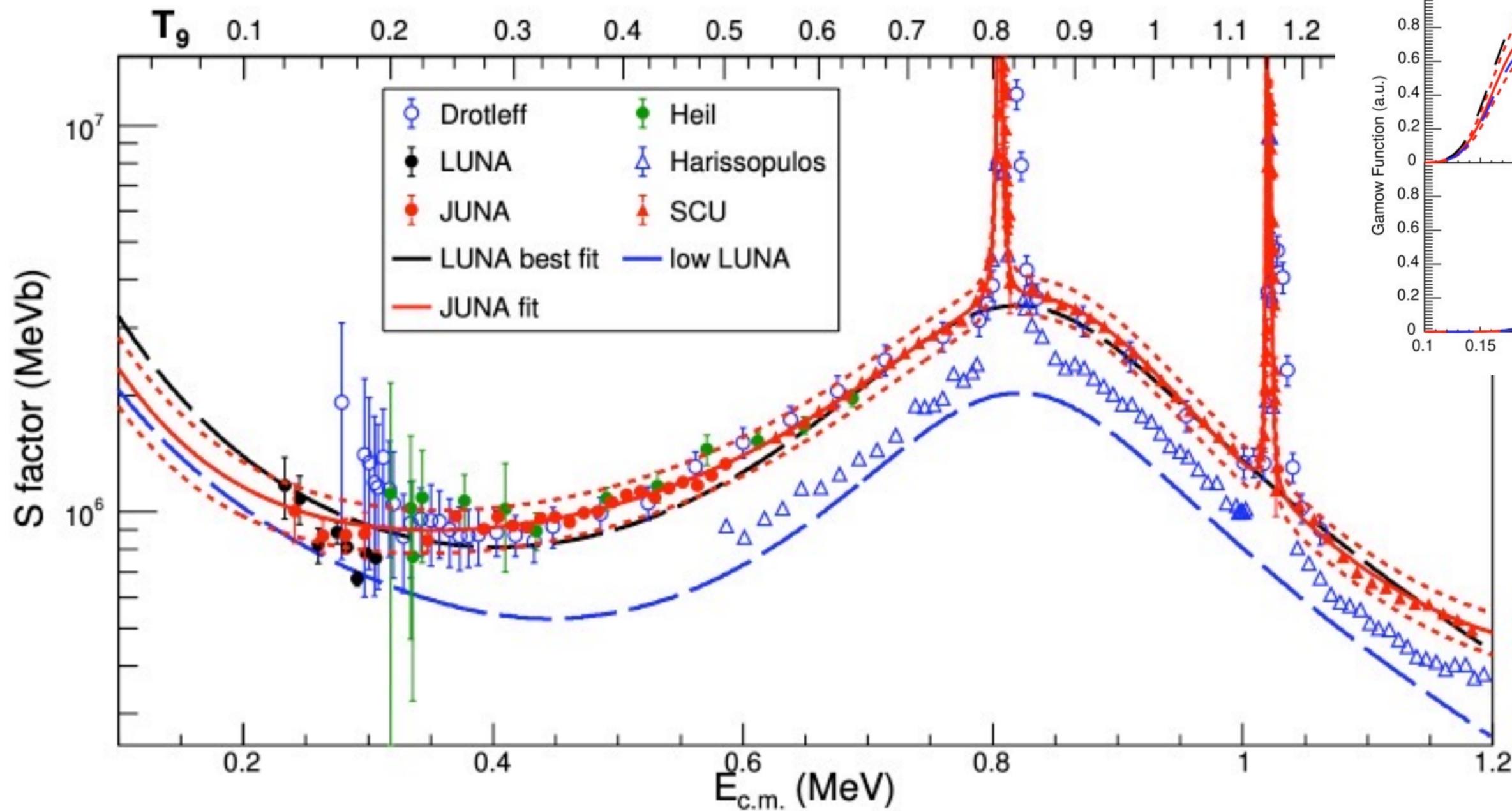
- Beam intensities from tens of μA up to 2.5 pA
- Use thick target to eliminate target-thickness related uncertainties
- He^{++} beam extends the beam energy up to 800 keV (Use slits to remove D^+)

$^{13}\text{C}(\alpha, n)^{16}\text{O}$ Measurement



JUNA2021

SCU2022



- mA thick target, differential method to pin down thickness
- magnetic removal of D⁺ from He²⁺, cover 0.4 MeV to 0.8 MeV (JUNA), cover i-process; to 1.9 MeV tandem, calibration of eff., cross check other data
- n background 5/hour, 2.5 MeV eff. 25%, good S/N

JUNA results from Run -1

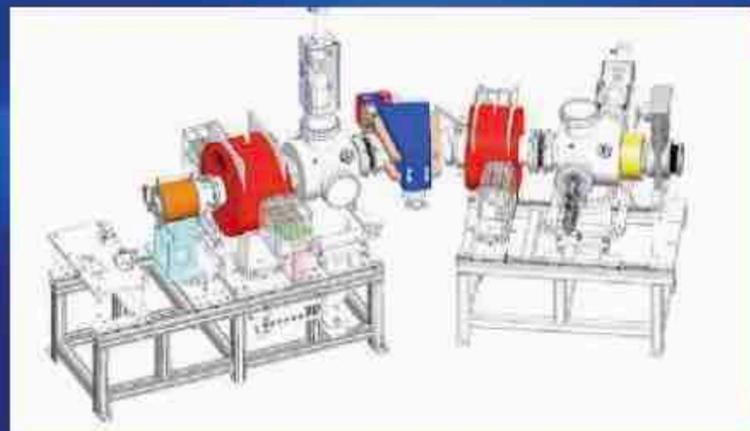


Reaction	Quantities	Best data before	JUNA data	Publication
Holy grail $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$	Lowest energy/keV	891	552	In preparation
	Cross section/b	10^{-11}	10^{-12}	
Neutron source $^{13}\text{C}(\alpha, n)^{16}\text{O}$	Energy range/keV	230-300	240-1900	PRL 129(2022)132701
	s-process	50%	20%	
^{26}Al abundance $^{25}\text{Mg}(p, \gamma)^{26}\text{Al}$	Uncertainty	21%	8%	Science Bulletin 67(2022)2 cover paper
F abundance $^{19}\text{F}(p, \alpha \gamma)^{16}\text{O}$	Lowest energy/keV	189	72	PRL 127(2021)152702 Editor suggestion
	Uncertainty	80%	5%	
Ne isotope ratio $^{18}\text{O}(\alpha, \gamma)^{22}\text{Ne}$	Uncertainty	472 ± 18 keV	474.1 ± 1.1 keV	PRL 130(2023)092701
CNO breakout $^{19}\text{F}(p, \gamma)^{20}\text{Ne}$	Lowest energy/keV	300	200	Nature 610(2022)656 news and views

Green lights for JUNA Run-2

Courtesy of W.P. Liu

- CJPL IAC highly recommend JUNA and gave green lights for next 5 years and support JUNA using A1 space
- High density radiation hard target and gas target, higher efficiency neutron and gamma detectors
- Run 2 proposal evaluated and approved from July 2025 to February 2026



Improved ion source



Run-2 kickoff meeting April 24, 2024, CIAE



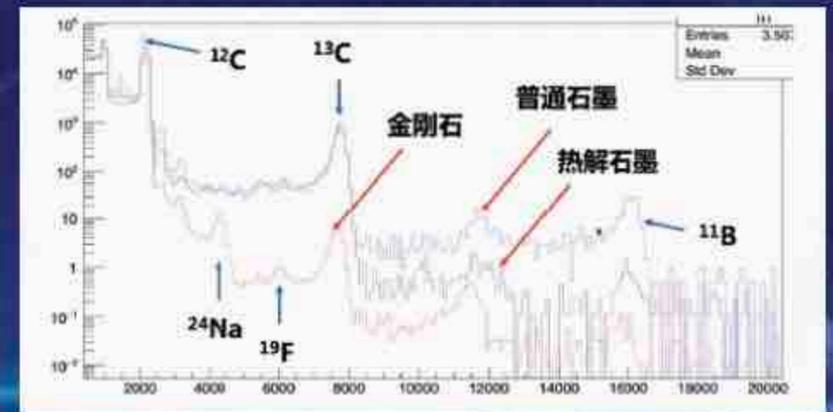
floor plan for JUNA Run-2



Gas jet target



Run-2 plan approved May 15, 2025, BNU



test result for diamond target



Enlarged BGO array



Sept., 2024
CJPL-II A1 ready



October, 2024
Accelerator in A1



May, 2025
Accelerator ready for beam

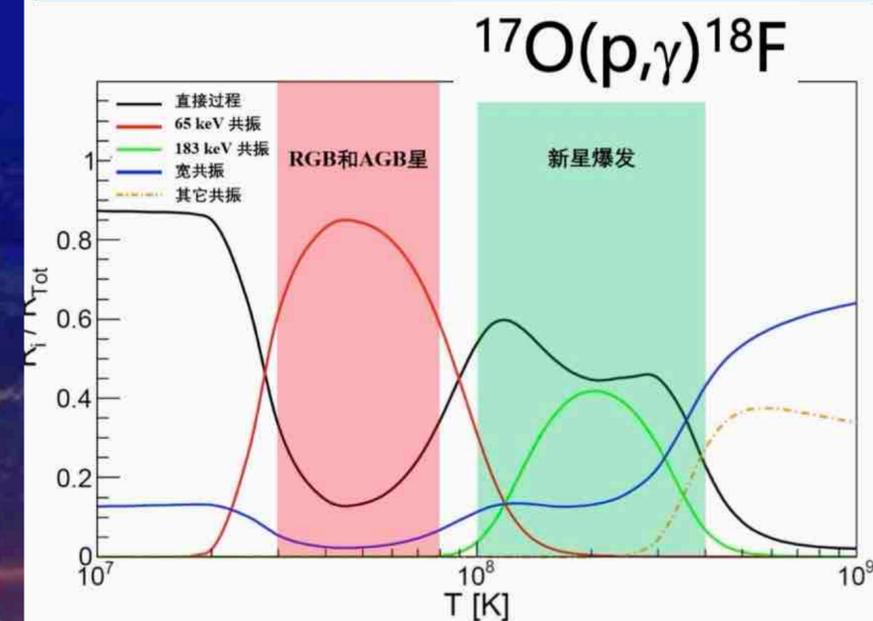
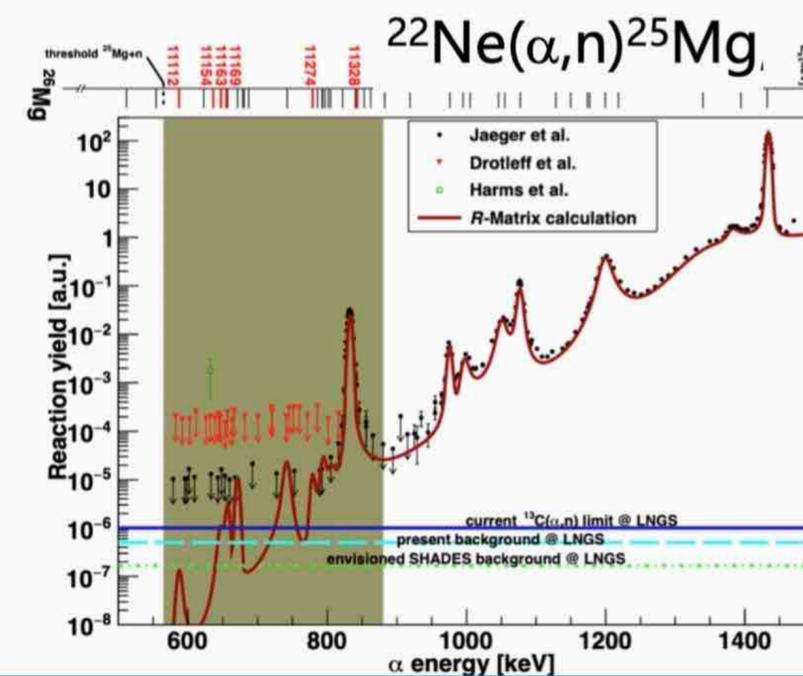
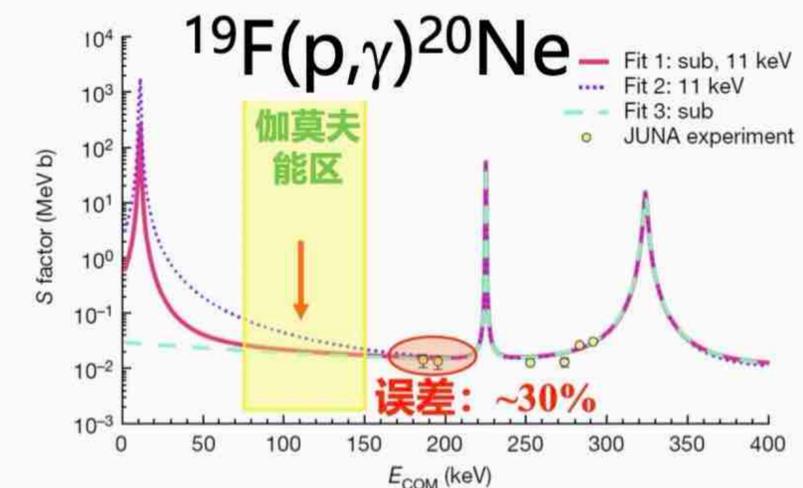
JUNA Run-2 Exp.: 2025-2027

From Run-1 to lower energy

- $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$, precision from 1s to 3s
- $^{13}\text{C}(\alpha, n)^{16}\text{O}$, full coverage of s-process
- $^{19}\text{F}(p, \gamma)^{20}\text{Ne}$, cover 80-150 keV with high precision
- $^{14}\text{N}(p, \gamma)^{15}\text{O}$, Solar neutrino

Using gas target

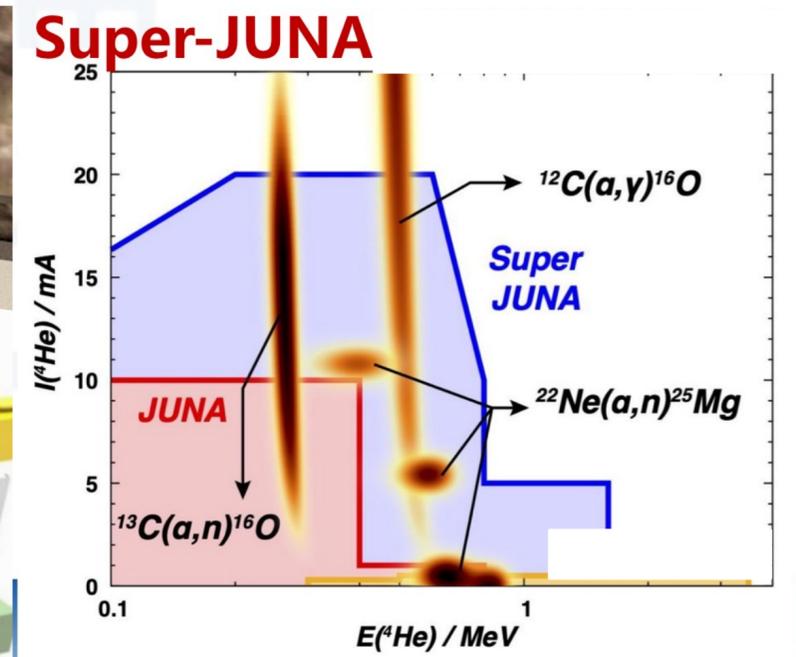
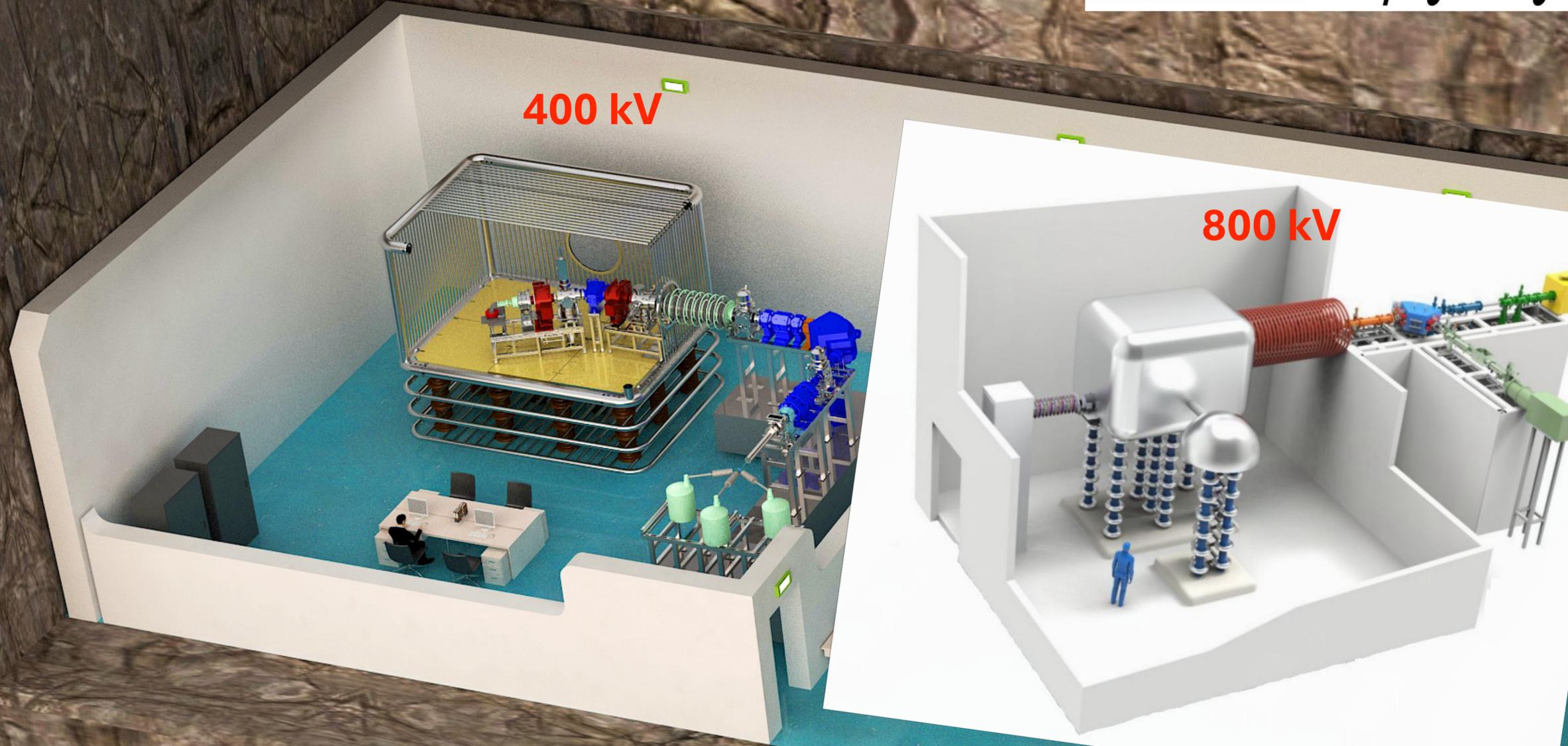
- $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$, weak s-process n source
- $^3\text{He}(\alpha, \gamma)^7\text{Be}$, solar neutron, Li problem, 80-380 keV





锦屏深地核天体物理实验

Jinping Underground Nuclear Astrophysics Experiment



JUNA

Super JUNA

2024-2030 JUNA and Super-JUNA



	JUNA	Super-JUNA	JUNA Exp.	Super-JUNA Exp.
2024	ground test run Run-2 exp.	R&D	(p,g)	
2025	Run-2	Ground test and fabrication	(a,n) and (a,g) gas target	
2026	Run-2	Setup and test	cont.	
2027	Upgrade Run-3 exp.	Test run Run-4 exp.	Test run	(p,g) test
2028	Run-3	Run-4	cont.	(a,g), (a,n) Exp.
2029	Run-3	Run-4	cont.	(a,g), (a,n) Exp.
2030	Run-3	Run-4	cont.	(a,g), (a,n) Exp.

JUNA and Super JUNA coverage

H burning



He burning



N source



C\O burning



γ -astronomy



JUNA achieved

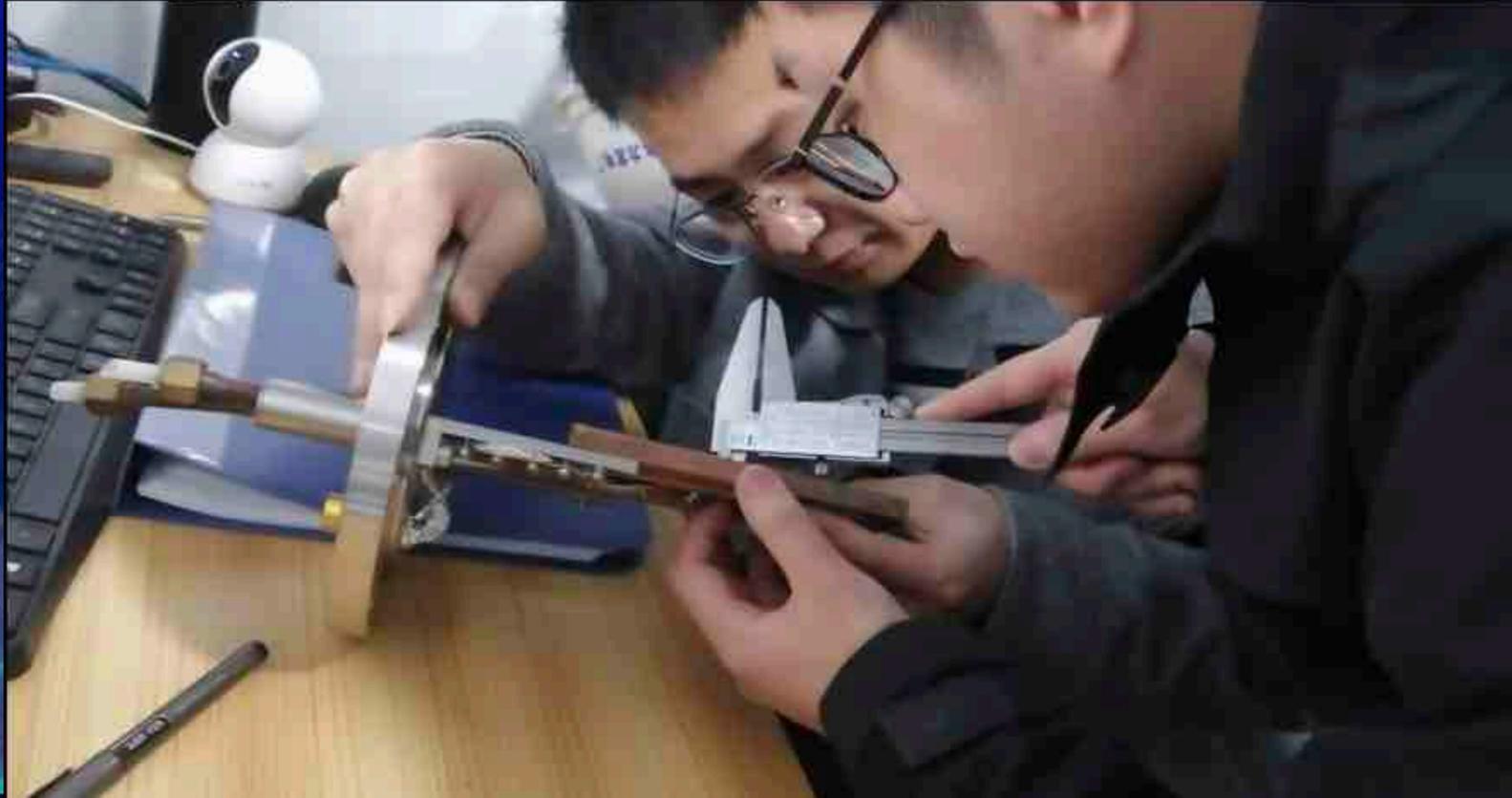


Super JUNA proposed

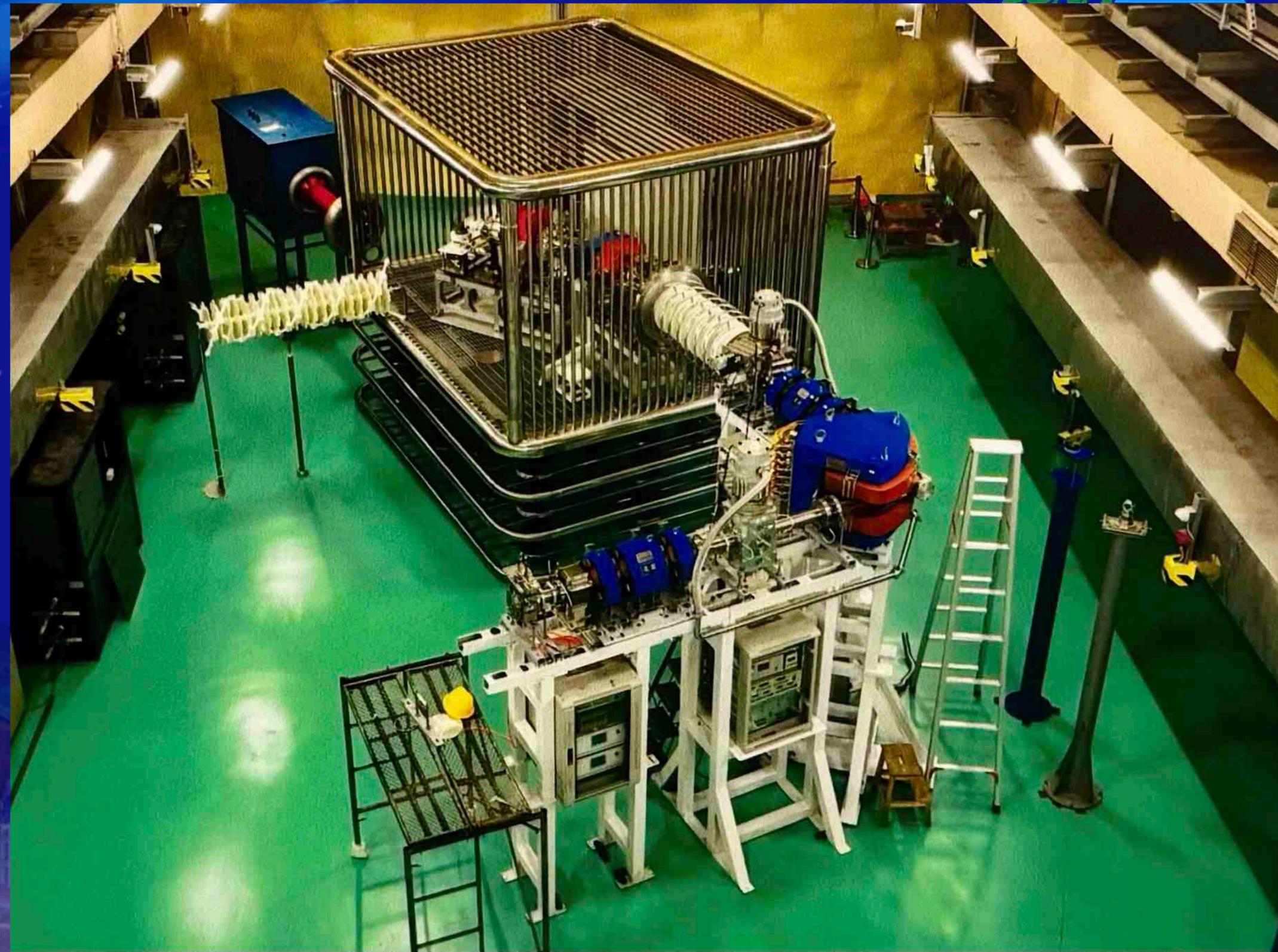
Summary



- **JUNA is an advanced deep astrophysics platform. China, follow Italy and United States and others, started to carry out direct measurement of key astrophysical reactions, which leading the nuclear astrophysics to the stage of precision numerical simulation stage**
- **JUNA accurately measured key nuclear astrophysical reactions, compared with previous experiment, beam intensity is higher, detector efficiency, target exposure, sensitivity and energy coverage are greatly improved**
- **From JUNA Run-1, Gamma-ray astronomical reaction has reached the highest precision, and the astrophysical holy grail reaction has achieved the highest sensitivity, new resonances revealing the origin of heavy element abundance in the oldest stars, and the discrepancies of neutron source reactions was resolved**
- **JUNA Run-2 started by the end of 2025, welcome to join JUNA collaboration and submit your proposals deep underground!**



Some more pictures





Köszönöm!

Grazie!

Thank you!



谢谢!

